

operator's finger he would be able to feel the pull of the earth's magnetic force upon it. The work done by a few light touches of the finger must be amply sufficient to furnish all the energy required to deflect the needle. But to a modern electrician it certainly seems a remarkably efficient transformation.

GEORGE J. BURCH.

An Introduction to the Study of Colour Phenomena. By Joseph W. Lovibond. Pp. 48; 10 coloured plates. (London: E. and F. N. Spon, Ltd.; New York: Spon and Chamberlain, 1905.) Price 5s. net.

The author states that his object has been to supply the long-felt want of a power of recovering a given colour sensation and of a colour nomenclature by which that sensation may be quantitatively described. To this end "scales of red, yellow and blue were constructed of glass slips, the slips of each scale being all of one colour with a regular variation in intensity from 0.01 to 20 units, equal units of the three scales being in colour equivalence with each other. . . . The test of equivalence is that a white light viewed through equal units of the three scales should give no evidence of colour. . . . The fogs on Salisbury Plain furnished the light actually used." It was found that red, yellow, and blue were the only colours suitable for systematic work, and that any colour could be produced by their combination. The dimensions of the unit are, it is said, necessarily arbitrary, but the scale-divisions are equal, while the unit itself is recoverable.

The colour to be tested is matched by that or the light transmitted by one of the glasses, or by several superposed, equality of luminosity being secured, when necessary, by the interposition of a neutral-tinted combination between the eye and the coloured object. A specification of the glasses employed is registered, according to certain rules, as a formula which defines in terms of the author's constants the colour "developed," and supplies data for its future reproduction.

To those who are accustomed to regard the spectrum as the natural basis of colour experiment the author's method cannot but appear crude and unscientific; but, given a sufficient supply of carefully selected glasses, it is probable that much useful work might be done in a rough and ready way by its means. An example occurs in the quantitative study of the colour of the human blood in health and in disease, which is illustrated in plate vi.

The book concludes with an exposition of Mr. Lovibond's new theory of colour.

Index Phytochemicus. By Drs. J. C. Ritsema and J. Sack. With introduction by Dr. M. Greshoff. Pp. 86. (Amsterdam: J. H. de Bussy.)

DR. GRESHOFF explains in the introduction to this volume that it originated in a card index to the literature of plant chemistry compiled for use in the laboratory of the Colonial Museum at Haarlem, where the work carried on consists principally of the investigation of the proximate constituents of plants.

The index enumerates the names of more than two thousand plant constituents, and gives in each case the percentage composition, formula, melting or boiling point, and at least one reference to the literature—usually Beilstein's "Handbuch," though in a few cases the references are to original papers. The volume also contains a short but useful bibliography of plant chemistry.

The information given in the tables, so far as can be judged from trials in a few cases, appears to be accurate, and the index should prove useful to chemists engaged in the investigation of plant products.

LETTERS TO THE EDITOR.

[The Editor does not hold himself responsible for opinions expressed by his correspondents. Neither can he undertake to return, or to correspond with the writers of, rejected manuscripts intended for this or any other part of NATURE. No notice is taken of anonymous communications.]

Eclipse Predictions.

It is always interesting to compare the results of observation with those predicted by calculation. In the case of the recent total eclipse of the sun this is rendered difficult by the want of agreement in the predictions of the two most used authorities, the Nautical Almanac and the *Connaissance des Temps*. The discrepancies in the predicted duration of totality and of the breadth of the band traced on the earth's surface by the total phase are made apparent in the following table. It is compiled from the table in the Nautical Almanac headed "Limits of total phase of the Solar Eclipse," and the corresponding table in the *Connaissance des Temps* entitled "Limites de l'Éclipse totale et Durée de la Phase totale sur la Ligne centrale." Entries for as nearly as possible the same time in each table have been taken and are placed together:—

Column A contains the authority, Nautical Almanac (N.A.) or *Connaissance des Temps* (C.T.).

Column B contains the time (G.M.T.) for which each prediction is made.

Column C contains the calculated distance (in nautical miles) and the bearing of the northern limit of totality from the corresponding southern limit.

Column D contains the durations of totality on the central line as predicted by the one authority and (in brackets) as interpolated from the prediction of the other.

Column E contains the differences of these pairs of values.

A	B	C	D	E
1905 Aug. 30	Distance	Bearing	N.A.	C.T.
G.M.T.			secs.	secs.
h. m.				
C.T. 0 22 ... 113.5 ... N. 0° W. ... (198 4) ... 206 ... 7.4				
N.A. 0 24 ... 101.5 ... , 2 W. ... 200.6 ... (208) ... 7.6				
C.T. 0 35.2 ... 109.5 ... , 2 E. ... (211) ... 219 ... 8.0				
N.A. 0 36 ... 102 ... , 11 ... 211.8 ... (219.5) ... 7.7				
C.T. 0 50.3 ... 114 ... , 6 ... (220.2) ... 228 ... 7.8				
N.A. 0 48 ... 104 ... , 19 ... 219.1 ... (227.4) ... 8.3				
C.T. 1 7.0 ... 116.5 ... , 10 ... (223 8) ... 231 ... 7.2				
N.A. 1 8.0 ... 104 ... , 31 ... 223.8 ... (231.2) ... 7.4				
N.A. 1 24 ... 105.5 ... , 37 ... 22.7 ... (226.6) ... 5.9				
C.T. 1 24.9 ... 116.5 ... , 12 ... (220.2) ... 227 ... 6.8				
C.T. 1 43.1 ... 115 ... , 14 ... (209.2) ... 215 ... 5.8				
N.A. 1 44 ... 106 ... , 44 ... 208.4 ... (214) ... 5.6				

It will be seen that, for stations in Spain and the adjacent Mediterranean, the duration of totality on the central line was predicted by the French authority to be from seven to eight seconds longer than by the British authority. In the same region, the width of the band of totality is from ten to eleven nautical miles greater by the French than by the British prediction. The orientation of the line connecting the two limits of totality also differs considerably in the two tables.

It is reported that at Sousse and Gabes, two towns in Tunisia, the eclipse was partial, while a total eclipse had been predicted for them. The prediction for these places would surely rest on French authority; we are therefore entitled to conclude that the mistake has been made by the French calculators. An excessive estimate of the width of the band of totality would almost certainly be accompanied by an excessive estimate of the duration of totality, and the table shows that both estimates are considerably greater in the *Connaissance des Temps* than in the Nautical Almanac.

J. Y. BUCHANAN.

October 13.

Absence of Vibration in a Turbine Steamship.

RETURNING homeward to Paris the middle of September from the Tripoli eclipse, and finding passage to America difficult to obtain, I chanced to learn that the triple-screw turbine steamer R.M.S. *Virginian* was sailing from Liverpool for Montreal on September 30, so I was very glad to have the opportunity of a voyage in a ship full powered with this novel type of propulsion. After a week on board I have no hesitation in saying that for freedom from

the nerve-annoying tremors incident to the usual reciprocating engines, the *Virginian* has proved far and away the quietest steamship I have ever voyaged on. Excellent evidence of this, I think, lies in the exceptionally large number of passengers who dined comfortably in the saloon at the roughest period of our entire passage. There was a fairly heavy sea on, and the ship was by no means free from wave-originated motion. So I am quite of the opinion that sea-sickness and all its train of discomforts must be greatly aggravated by the engine-borne tremors of the ordinary steamship, and that many people who are delicate sailors under ordinary conditions might take ocean journeys with comparative comfort in a turbinised ship.

So unostentatious are the rotary engines of the *Virginian*, let alone their occupying but one-fourth the space of the usual expansion engines, that the quietness of their powerful and effective working, in every part of the ship, was continually deceiving one into thinking that the vessel had lost headway, or might have come to anchor altogether. Especially was this true in the dining saloon, that most critical of all spots, where one could rarely detect so much as a ripple on water in a glass, although going ahead at full speed of 15 knots.

To my mind the *Virginian* seemed to behave all the voyage quite as if her motive power were entirely without her; in fact, she could scarcely have ridden more smoothly, or with less of that exasperating vibration (the unceasing action of which, I am convinced, is a prominent factor in inducing *mal de mer*), if she had been towed at the identical speed by a huge hawser.

DAVID TODD.

R.M.S. *Virginian*, Straits of Belle Isle, October 4.

A Parasite of the House-fly.

REVERTING to the recent correspondence under this heading between Mr. Davenport Hill and Prof. Hickson (*NATURE*, August 24 and 31), I recall that a few years back many house-flies with Chelifers attached were sent to me at the Natural History Museum for determination of the species and explanation of the phenomenon. The first task was as easy as the second was difficult. The Chelifer was in most, nay in all, cases, so far as my memory serves, *Chernes nodosus*. But those who suggest that the explanation is to be sought and found in the value of the habit as a means of securing dispersal hardly realise, I think, the difficulties in the way of its acceptance. Chelifers are minute, active, and, for arthropods, not exceptionally prolific. Hence the sufficiency of "elbow-room" for the survivors of a family of, say, forty, on the site chosen by the female for her progeny does not coincide with the view that they have special need of transportation. Moreover, when we remember that a Chelifer attached to a fly is exposed to the danger of being killed by the enemies of that insect, and also to the great chance of being landed in a wholly unsuitable environment, it can hardly be maintained that the advantage derived from this method of dispersal has been a sufficiently important factor in survival to preserve and foster an initial instinct to grab and hang on to the legs of flies. That the aerial portage thus secured, whether fortuitously or "intentionally," must be a means of dispersal is too obvious to dispute; but I do not think more than that can be claimed for it, since it is as likely to end in failure as in success.

Chelifers may be found not uncommonly beneath the wing-cases of large beetles. Presumably this habitat has been adopted for the sake of the food supplied by the parasitic mites infesting the beetles. This fact, I think, suggests a line of investigation which may lead to a more satisfactory explanation of the association between Chelifers and flies than that put forward in Prof. Hickson's letter.

Zoological Gardens, October 14.

R. I. Pocock.

Incandescence of Meteors.

It is with great diffidence that I approach this difficult subject, but the theory that the incandescence of meteors is due to the heat generated by the friction between these bodies and the molecules of gas composing our atmosphere

I have always found difficult to believe. The following theory is one which has occurred to me, and seems quite a plausible one. Meteors are usually of a metalliferous nature, and consequently will have a comparatively low electrical resistance. When they approach the earth they will enter a magnetic field, and they will cut the lines of force of this field at a high velocity. A high electrical potential will be generated, and consequently electric currents which will be inversely proportional to the resistance. The electrical energy thus produced will be dissipated in heat, and if of sufficient intensity will raise the meteor to incandescence. The truth or otherwise of this theory could, I believe, be calculated, as the data necessary for doing so will be at the disposal of readers of *NATURE* who make this branch of astronomy their study. This theory may have already been advanced, as I am not in touch with the latest developments of the science.

Coatbridge, September 5.

GEORGE A. BROWN.

THE electric currents which the author of the above letter regards as possibly constituting an efficient source of the luminosity of meteors must no doubt arise, and play a certain part in the heat and light development. But the measure in which they can be supposed to contribute to it must clearly be extremely small; or rather, it must be incomparably subordinate to the intense ignition of the air produced, not at all by friction,¹ but by the air's adiabatic compression against the front surface of the meteorite; which is certainly quite competent, by itself alone, to develop what may be said to approach pretty nearly to fabulous degrees of temperature. If the kinetic energy of translation, in foot-pounds ($v^2/2g$), of 1 lb. of the air propelled (at, say, 30 miles per second) with the meteor's speed (v feet/sec.) on its front face, be divided by 330, the number thus obtained ($1,180,620^\circ$ C., in the case supposed) will be the number of centigrade degrees through which it will be heated by the pure process of compression, supposing that the air can continue to subsist at all with its ordinary mechanical deportment and thermodynamical properties unaffected at that enormously high temperature. In the further forward, gradually advancing layers, and in the laterally escaping currents of the air, on which the high forward speed of the meteor is only partially impressed, and which move more slowly on their various courses, the compressions are correspondingly less, and the lower but still exceedingly high temperatures can be similarly calculated from any fair estimates of the air's collective or absolute velocity of translation in those different positions.

It is in the different rates of transport of these heated air-streams, all of them, as well as the highly attenuated motionless atmosphere around, affording very easy passages to electricity, across the earth's magnetic field or system of lines of magnetic force, that fitting circuits can certainly be found (either passing through, or else entirely omitting the meteorite itself), in which, in the way suggested in the above letter, electric currents may be quite certainly concluded to be magneto-electrically induced. For while one part of a closed air-circuit resting against the meteorite's front surface, and another part of it situated in the still atmosphere in front of or behind it, would be journeying towards or from each other with full meteor-speed, the circuits so composed would be most suitably conditioned for developing induced currents round them by

¹ Although a very general belief, it is as yet an entirely mistaken supposition that the high speed of impact of a meteorite into the rarer regions of the atmosphere reduces the air, by giving it no time to dissipate itself in front of the meteorite, to a state of granulation, or to a wedged throng of molecules producing heat by friction *inter se* and against the surface of the meteorite. Just the reverse of this condition is, however, really true, that the air remains a perfectly and frictionlessly elastic fluid, however much it is compressed and intensely heated by the impact. The speeds of sound-waves in the heated air which perform the office of transmitting and maintaining the orderly array of pressures in the streaming flows, at length differ in defect, in fact, from the air's speeds themselves in proportions which, as those mount up to meteor-speeds of many miles per second, only decline asymptotically to about the ratio $1:\sqrt{5}$, or nearly $1:2\frac{1}{2}$. Since, then, these sound-waves, which convey the strokes and shocks of the collision to and fro between the meteor-centre and the surrounding air, arise and travel in the moving field of the compressed air as if it were at rest, it is easy to perceive that by their extremely rapid actions a most exceptionally perfect elastic-fluid relation, or steady disposition of the lines, or lanes of air-flow and blast-pressure, must really be established and maintained in evenly persistent shapes and contour, in the swirl of incandescent air which forms the meteor's head.